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1Multifractal comparison of the outputs of two optical disdrometers

2

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6

7Abstract:

8 In this paper a Universal Multifractals comparison of the outputs of two types of
9collocated optical disdrometers installed on the roof of the Ecole des Ponts ParisTech is
10performed. A Campbell Scientific PWS100 which analyses the light scattered by the
11hydrometeors and an OTT Parsivel² which analyses the portion of occluded light are
12deployed. Both devices provide the binned distribution of drops according to their size and
13velocity. Various fields are studied across scales: rain rate (R), liquid water content (ρ),
14polarimetric weather radars quantities such the horizontal reflectivity (Z_h) and the specific
15differential phase (K_{dp}), and DSD parameters such as the total drop concentration (N_t) and the
16mass-weighted diameter (D_m).

17 For both devices a good scaling is retrieved on the whole range of available scales (2h
18– 30s), except for the DSD parameters for which the scaling only holds down to few minutes.
19For R , the UM parameters are found equal to 1.5 and 0.2 for respectively α and C_1 . Results
20are interpreted with the help of the classical $Z_h - R$ and $R - K_{dp}$ radar relations.

21

22Key words: multifractals, disdrometer, radar

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25

261) Introduction

27

28 Rainfall measurement with the help of disdrometers is rapidly developing for point
29 measurements. The first ones were impact disdrometers (Joss and Waldvogel 1967), and now
30 optical ones (Löffler-Mang and Joss 2000, Ellis et al. 2006, Battaglia et al. 2010, Frasson et
31 al. 2011) are more commonly used for operational and research purposes. The great advantage
32 of these devices with regards to more conventional rain gauges is that they do not measure
33 only a rain rate (or rainfall depth) but information about size and fall velocity for all the
34 hydrometeors passing through the sampling area whose size is few tens of cm^2 . The Drop Size
35 Distribution (DSD) can then be computed from this raw data. From the DSD it is possible to
36 estimate numerous rain related fields such as the rain rate (R) or the liquid water content (ρ)
37 or even quantities measured directly by polarimetric weather radars such as the horizontal
38 reflectivity (Z_h) and the specific differential phase (K_{dp}) (Jaffrain and Berne 2012, Leinonen et
39 al. 2012, Verrier et al. 2013). This widens a lot the range of applications for disdrometer data.

40 Numerous experiments have been carried out to compare the rain rate output of
41 various types of collocated disdrometers along with rain gauges (Miriovsky et al. 2004,
42 Krajewski et al. 2006, Frasson et al., 2011, Thurai et al. 2011). The temporal evolution
43 (Thurai et al 2011) or spatial distribution (Jaffrain and Berne 2011) of total drop concentration
44 (N_t) and mass-weighted diameter (D_m), which are commonly used to fully characterize drop
45 size distribution, have also been analysed. However these comparisons are usually done only
46 at a single resolution, most commonly the maximum one available (i.e. using the time series
47 at the recording time step). In this paper we suggest to carry out the comparison not only at a
48 single scale but across scales. Considering various scales at once enables to obtain results
49 robust over a wide range of scales. To achieve this, a theoretical framework relying on scale
50 invariant properties is implemented. Available data enables to carry out analysis on scales

51ranging from 30 s to approximately 2 hours for a punctual measure. Moreover not only R , ρ ,
52or N_t and D_m , but also the radar quantities Z_h and K_{dp} will be investigated. The output data of
53two disdrometers, which have been deployed for few months on the roof of the Ecole des
54Ponts ParisTech building, will be used for this study. The two disdrometers are a Campbell
55Scientific PWS100 (Ellis et al. 2006, Campbell Scientific Ltd 2012), which has seldom been
56used in such study, and an OTT Parsivel² (Battaglia et al. 2010, OTT 2014).

57 The variability across scales will be quantified with the help of stochastic universal
58multifractals which have been extensively used to analyze, model, and simulate geophysical
59fields extremely variable over wide range of scales such as rainfall (Schertzer and Lovejoy
601987, 1997, Marsan et al. 1996, Olsson and Niemczynowicz 1996, Harris et al. 1997, de Lima
61and Grasman 1999, Lovejoy and Schertzer 2007, Nykanen 2008, Royer et al. 2008, de Lima
62and de Lima 2009, Mandapaka et al. 2009, de Montera et al. 2009, Verrier et al. 2010). This
63framework has seldom been applied to Z_h (see Tessier et al. 1993 for an example in space and
64not in time as here), and not to K_{dp} to the knowledge of the authors.

65 Data and the retrieval of the various studied fields are presented in section 2 along
66with a brief reminder of the Universal Multifractals (UM) framework and a presentation of
67the methodology implemented. Results are in section 4 where classical $Z_h - R$ or $R - K_{dp}$
68relations are investigated with this data set, and scaling features and UM parameters estimates
69of the various fields are discussed.

70

712) Data and methods

722.1) Description of the disdrometer data

73 The data used in this paper was collected between 27 September 2013 and 18 January
742014 by two disdrometers installed on the roof of the Ecole des Ponts ParisTech building.
75They are both made of a transmitter that generates one or several laser sheet(s) and

76 receiver(s). The Campbell Scientific PWS100 computes size and fall velocity from light
77 refracted by the hydrometeors (Ellis et al. 2006, Campbell Scientific Ltd 2012) whereas the
78 OTT Parsivel² relies on occluded light (Löffler-Mang and Joss 2000, for an initial version;
79 Battaglia et al. 2010, OTT 2014). The 50 recorded events of this period are used in this study.
80 The criteria defining an event is a rainy period during which more than 1 mm is collected and
81 that is separated by more than 15 min of dry conditions before and after. The main output of
82 the disdrometers is a matrix with the number n_{ij} of drop recorded according to classes of
83 equivolumic diameter (index i , and defined by a centre D_i and a width ΔD_i expressed in mm)
84 and terminal fall velocity (index j , and defined by a centre v_j and a width Δv_j expressed in
85 $\text{m}\cdot\text{s}^{-1}$). This matrix is recorded for each 30s time step (Δt). Gires et al. (2014) which used the
86 same data set noticed that the oblateness of drop was not properly taken into account in the
87 PWS100 rationale and suggested a correction which is used here. Furthermore, as suggested
88 by various authors (Jaffrain and Berne 2011, Kruger and Krajewski 2002, Thurai and Bringi,
89 2005) all the drops whose velocity was more than 60% different from what was expected by
90 Beard's model (Beard 1977) according to its size were removed because considered as non
91 meteorological measurements.

92

93 The rain rate for each time step is then computed as:

$$94 R_{PWS/Pars} = \frac{\pi}{6\Delta t} \sum_{i,j} \frac{n_{i,j} D_{PWS/Pars,i}^3}{S_{eff}(D_i)} \quad (1)$$

95 where $S_{eff}(D_i)$ is the sampling area of the device which is slightly modified

96 according to the drop size to take into account potential edge effects for large drops. For the

97 Parsivel² we have $S_{eff}(D_i) = L \left(W - \frac{D_i}{2} \right)$, where $L = 180$ mm and $W = 30$ mm are respectively

98 the length and width of the sensing area ($LW = 54$ cm^2) (OTT 2014). The PWS100 is not

99subject to this issue and S_{eff} is taken as constant equal to 40 cm² (Campbell Scientific Ltd
1002012). For the selected events the total rainfall depth collected is equal to 172 mm for the
101PWS100 and 154 mm for the Parsivel². A tipping bucket rain gauge is also located on the
102same roof and collected roughly 170 mm which is more in agreement with the PWS100, but it
103should not be “over-interpreted” given that its accuracy is not very high due to high rain rates
104and low level of maintenance at the beginning of the recording period. The normalized bias
105between the two disdrometers (computed for the time steps where $R > 1\text{mm/h}$) is equal to
1060.11; the correlation equals 0.96; the Nash-Sutcliffe efficiency coefficient equals 0.89 and the
107RMSE is 1.83. Although not negligible the differences between the two disdrometers are
108rather low compared to what is commonly observed (Miriovsky et al. 2004, Krajewski et al.
1092006, Frasson et al., 2011, Thurai et al. 2011).

110 A discrete drop size distribution (DSD) is computed from the available data as:

$$111 N(D_i) = \frac{1}{S_{eff}(D_i)\Delta D_i\Delta t} \sum_j \frac{n_{i,j}}{v_j} \quad (2)$$

112The number of drops with a diameter in the class i per unit volume (in m⁻³) is given by
113 $N(D_i) \Delta D_i$. As it is commonly done, we use the total drop concentration N_t (m⁻³) and mass-
114weighted diameter D_m (mm) to characterize the DSD. In this framework the DSD is written as
115 $N(D) = N_t f(D_m)$; where it appears that D_m characterizes the shape of the DSD and N_t its
116total intensity. These two parameters are defined as (Jaffrain and Berne 2012a, Leinonen et al.
1172012):

$$118 N_t = \int_{D_{min}}^{D_{max}} N(D)dD \quad (3)$$

$$119 D_m = \frac{\int_{D_{min}}^{D_{max}} N(D)D^4 dD}{\int_{D_{min}}^{D_{max}} N(D)D^3 dD} \quad (4)$$

120 For the practical computation of these quantities (and the other ones), the integral is replaced
 121 by a summation over all the classes of diameter, i.e.:

$$122 \int_{D_{\min}}^{D_{\max}} f(D)N(D)dD \approx \sum_i f(D_i)N(D_i)\Delta D_i \quad (5)$$

123 Lastly the liquid water content (ρ in g.m^{-3}) the horizontal reflectivity (Z_h in $\text{mm}^6.\text{m}^{-3}$)
 124 and the specific differential phase (K_{dp} in $^{\circ}.\text{km}^{-1}$) are estimated with the help of the DSD from
 125 which they are basically various moments. These physical parameters are given by Jaffrain
 126 and Berne (2012a) and Leinonen et al. (2012):

$$127 \rho_l = \rho_w \frac{1}{10^3} \frac{4}{3} \pi \int_{D_{\min}}^{D_{\max}} N(D) \left(\frac{D}{2} \right)^3 dD \quad (6)$$

128 With ρ_w , the liquid water density in g.m^{-3} ,

$$129 Z_{h,v} = \frac{\lambda^4}{\pi^5 \left| \frac{m^2 - 1}{m^2 + 1} \right|} \int_{D_{\min}}^{D_{\max}} N(D) \sigma_{B;h}(D) dD \quad (7)$$

$$130 K_{dp} = \frac{10^{-3} 1800 \lambda}{\pi} \int_{D_{\min}}^{D_{\max}} \text{Re}[S_{hh}(D) - S_{vv}(D)] N(D) dD \quad (8)$$

131 Where $\sigma_{B;h}$ (in mm^2) is the backscattering cross section at horizontal polarization, and
 132 $\text{Re}(S_{hh/vv})$ (in mm) is the real part of the forward scattering amplitude at horizontal/vertical
 133 polarization, λ is the radar wavelength (in mm) and m the complex refractive index of water.

134 The scattering coefficients were computed with the help of the Python PyTMatrix library
 135 (Leinonen 2014) which relies on the T-Matrix code by Mishchenko et al. (1996).

136 Computations were carried out for a radar wave length equal to 53.5 mm corresponding to C-
 137 band radars (this wave length was chosen to facilitate comparison with other studies, because
 138 it is the most widely used and corresponds to the radar currently mostly used by Western

139 European meteorological services), a temperature of 20°C ($m=8.633 + 1.289i$) and an oblate
 140 spheroids model for drop shape with an axis ratio - equivolumic diameter relation

141 corresponding to the one implemented in the Parsivel² rationale (Battaglia et al. 2010). The set

142up considered for drop orientation was the same as in Leinonen (2012); i.e. drops are partially
143aligned and a normal distribution (mean and standard deviation respectively equal to 0° and
144 7° , in agreement with the findings of Bringi et al. 2008) characterizes the angle of the
145symmetry axis.

146

1472.2) Methodology: UM framework

148

149 The key elements of the theoretical framework of UM are presented here and the
150reader is invited to refer to Schetzer and Lovejoy (2011) for a recent review. In the following
151 ϵ_λ denotes a conservative field at resolution λ , defined as the ratio between the observation
152scale l and the outer scale L ($\lambda=L/l$). Practically ϵ_λ is obtained simply by up-scaling (averaging
153consecutive time steps) the measured field at the maximum resolution. If ϵ_λ is a multifractal
154field, then its statistical moment orders scale with resolution as:

$$155 \langle \epsilon_\lambda^q \rangle \approx \lambda^{K(q)} \quad (9)$$

156Where $K(q)$ is the moment scaling function that fully characterizes the variability across
157scales of the field ϵ_λ . The quality of the scaling is investigated with the help of the Trace
158Moment (TM) analysis which simply consists in plotting equation 9 in log-log, the slope of
159the obtained straight line being $K(q)$.

160 Most multiplicative processes converge toward Universal Multifractal which are fully
161characterized with the help of only two scale invariant parameters C_1 and α (this a broad
162generalization of the central limit theorem, Schetzer and Lovejoy 1987, 1997).
163 C_1 is the mean intermittency co-dimension and measures the clustering of the (average)
164intensity at smaller and smaller scales ($C_1=0$ for a homogeneous field). α is the multifractality

165 index ($0 \leq \alpha \leq 1$) and measures the clustering variability with regards to intensity level. In
 166 this specific framework which is implemented here $K(q)$ is given by:

$$167 K(q) = \frac{C_1}{\alpha - 1} (q^\alpha - q) \quad (10)$$

168 The UM parameter are estimated in this paper with the help of the Double Trace
 169 Moment (DTM) technique (Lavallée et al. 1993).

170 A common framework to deal with a non-conservative field ϕ_λ (i.e. we have $\langle \phi_\lambda \rangle \neq 1$)
 171 is to assume it can be written as

$$172 \phi_\lambda = \varepsilon_\lambda \lambda^{-H} \quad (11)$$

173 where H is the non-conservation parameter ($H=0$ for conservative fields), and ε_λ a
 174 conservative field characterized by a moment scaling function $K_c(q)$ depending only on UM
 175 parameters C_1 and α . More physically, this non-conservativeness means that the studied field
 176 ϕ_λ (the observations here) exhibits stronger correlations than the ones obtained with the help
 177 of a simple multifractal cascade and an additional fractional integration is needed to represent
 178 it. The moment scaling function $K(q)$ of ϕ_λ is given by:

$$179 K(q) = K_c(q) - Hq \quad (12)$$

180 H can be estimated with the help of (Tessier et al. 1993):

$$181 \beta = 1 + 2H - K_c(2) \quad (13)$$

182 where β is the spectral slope. It is the exponent of the power law that characterizes over large
 183 range of wave numbers the power spectrum of a scaling field:

$$184 E(k) \propto k^{-\beta} \quad (14)$$

185 Before going on let us clarify the relations between the notions of stationarity and
 186 conservation. For stochastic processes, stationarity refers to the fact that given statistics are

187invariant with respect to time translations. When no given statistics are mentioned, it usually
188corresponds to the strongest case of stationarity, i.e. the probability itself is time translation
189invariant and therefore all the statistical moments are stationnary The classical case of
190“second order stationarity” is presumably the most cited case and is particularly important for
191additive processes. It corresponds to time translation invariance of the second order moments
192of the increments. For multiplicative processes, conservation means that a given statistic is
193strictly independent from scale, in general this statistic is the mean of the field ($H=0$ in Eq.
19411). The latter implies a given form of stationarity because at all scales the fluctuations remain
195around this mean. However, these fluctuations are not only easily wilder than for an additive
196process, but their amplitude generally increases with smaller and smaller scales. Conversely,
197non-conservation implies a strong non stationarity, starting with the fact that the mean has a
198scaling behaviour (Eq. 11). For instance, a spectral analysis will detect the non conservation (
199 $H \neq 0$), due to the departure of the spectral slope, which is a second order statistic, from that
200of “a pink noise” (k^{-1}), where the exponent one corresponds here to the dimension of the
201embedding space (time series are studied here). With multifractal fields there is a further
202correction with the $K_c(2)$ (Eq. 13) which is not negligible (in the range 0.1-0.4 for the
203retrieved UM parameters).

204 The TM and DTM techniques are designed for analysing conservative fields ($H=0$)
205and remain reliable as long as $H<0.5$. In case of greater H , they should be implemented not on
206 ϕ_λ , but on the underlying conservative field ε_λ . A fractional integration of order H (equivalent
207to a multiplication by k^H in the Fourier space) is theoretically required to estimate ε_λ from ϕ_λ .
208However a common approximation (Lavallée et al. 1993) which provides reliable results
209consists in taking ε_λ at the maximum resolution simply equal to the renormalized absolute
210value of the fluctuations of the field i.e.:

$$211 \varepsilon_{\lambda} = \frac{|\phi_{\lambda}(i+1) - \phi_{\lambda}(i)|}{\langle |\phi_{\lambda}(i+1) - \phi_{\lambda}(i)| \rangle} \quad (15)$$

212 and then upscaling this field at other resolutions λ .

213 Spectral and multifractal analyses are implemented on ensemble average over various
 214 samples, i.e. each sample is considered as a realization of the process. For example it means
 215 that each sample is up-scaled and taken to the power q independently before taking the
 216 average in equation 9. The studied samples are extracted in the following way from the
 217 selected events: for each event (i) a sample size is chosen (necessarily a power of two); (ii) the
 218 maximum number of samples for this event is computed; (iii) the portion of the event of
 219 length equal to the sample size multiplied by the number of samples found in (ii) with the
 220 greatest cumulative depth is extracted; (iv) the extracted series is cut into various samples.

221 Given that the sample size is a power of two there is obviously some data lost during
 222 the selection process. Here the percentage of available data actually used is equal to 36, 65, 82
 223 and 91% for samples of size 512, 256, 128 and 64 times steps of 30s respectively. The chosen
 224 sample size should result from a trade off between the width of the available range of scales
 225 which should be as great as possible, and the amount of wasted data which reduces the
 226 reliability of the estimates. Here we chose to analyse the scaling properties with samples of
 227 size 256 (approximately 2h, 46 samples available), and given that no scaling break was
 228 identified on the main studied fields, UM parameters were estimated with samples of size of
 229 64 (approximately 30 min, 259 samples available) to benefit from the use of more data.
 230 Finally it should be mentioned that since D_m is not defined when there are no drops, only the
 231 samples containing drops at all time steps are used for this field.

232

233 3) Results and discussion

234 3.1) Standard radar relationships

235 Given that it was possible to retrieve both rain rates and radar parameters, it is possible
236 to analyse the validity, for this data set, of the standard power law relations that are commonly
237 assumed between these quantities:

$$238 Z_h = aR^b \quad (16)$$

$$239 R = cK_{dp}^d \quad (17)$$

240 These relations will be used in the following sections to help in the interpretation of the
241 observed scaling features exhibited by each quantity. Various authors (Campos and Zawadzki
242 2000, Jaffrain and Berne 2012b, Verrier et al. 2013) noticed a strong sensitivity of the
243 estimates of a and b on the method implemented to compute them. Here we performed an
244 orthogonal linear regression, which does not assume any dependent variable, on the logs of
245 Z_h , R and K_{dp} . The regressions are performed only for the time steps for which $K_{dp} > 10^{-2}$
246 because we noticed that small values in which we are not interested had a strong influence on
247 the retrieved parameters. Figures 1.a and 1.b display the regressions for the two relations and
248 both devices, and the retrieved values for a , b , c and d are shown in Table 1. It appears that the
249 quality of the fitting is comparable for the two relationships and slightly better for the
250 PWS100 than for the Parsivel² ($r^2 \sim 0.8$ vs. 0.75). The estimates are similar for both devices.
251 More precisely for the $Z_h - R$ relation a is slightly greater and b slightly smaller for the
252 PWS100 than for the Parsivel², meaning that the effects of each other are compensating. For
253 the $R - K_{dp}$ relationship c and d are both slightly greater for the PWS100, meaning a given
254 value of K_{dp} will systematically yield greater estimates of R with the PWS100 values. The
255 estimates of the exponents are in the range of those commonly observed (Jaffrain and Berne
256 2012b, Figuras I ventura et al. 2013, Verrier et al. 2013). With regards to a values they are in
257 agreement with those found by Verrier et al. (2013), and in the upper range of those reported
258 by Jaffrain and Berne (2012b). The c values are similar to those found by Jaffrain and Berne
259 (2012b). Relations 16 and 17 are studied only at the maximum resolution (30 s) and the strong

260scale dependency of the parameters a , b , c and d is not investigated here (see Verrier et al.
2612013 for an analysis of this issue for $Z_h - R$ relation). Parameters are computed here taking
262into account all the events at once, meaning that “climatic” (keeping in mind only 4 months of
263data are used) estimates are studied. An event based analysis will be discussed in future
264works. It is indeed not needed for the purpose of this paper, which is to quantify the scaling
265variability observed by the two disdrometers on various fields.

266 Finally the influence of using the retrieved “climatic” parameters for computing rain
267rates from radar parameters is assessed. With this purpose, the rain rates computed either
268directly from raw data (R) or through the radar relations 16 and 17 (respectively R_{Z-R} and
269 $R_{R-K_{dp}}$) once Z_h and K_{dp} have been estimated with raw data are compared. Figure 1.c and 1.d
270display, for the PWS100 data, a scatter plot for the $Z_h - R$ and $R - K_{dp}$ relations
271respectively. Similar curves are obtained for the Parsivel² data and not shown here. The
272scattering around the bisector is not negligible with a 30 s time step, and is more pronounced
273for the $Z_h - R$ than for the $R - K_{dp}$ relation which means that the use of “climatic” values is
274less acceptable for the former. It should also be noted that there is a tendency of
275underestimating large rain rate with the $R - K_{dp}$ relation, indeed $R_{R-K_{dp}}$ is systematically
276smaller than R for $R > 50$ mm/h. The total rainfall depths between the three techniques are very
277similar with roughly 3% or less differences.

278

2793.2) Scaling behaviour

280

281 The scaling features of the various fields are studied in this section. Figure 2.a displays
282the spectral analysis (i.e. equation 14 in log-log plot) of the rain rate for the PWS100. A very
283good scaling (i.e. a straight line) is observed on the whole range of available scales (30s – 2h).

284The spectral slope is greater than the embedding dimension of the field (1 for time series)
285suggesting that R is non-conservative (this will be confirmed in the next session with the
286estimation of H). A practical consequence is that the TM analysis which assumes a
287conservative field should not be implemented on the field itself but only on its conservative
288part which can be approximated by the absolute value of its fluctuations (equation 15). The
289results are displayed on Fig. 2.b, where the good scaling with a unique regime is confirmed.
290The r^2 for $q = 1.5$ which is taken as an indication of the quality of the scaling is greater than
2910.99. The same analysis carried out directly on the field yields non aligned points with a
292flattening for small scales ($r^2 = 0.94$ for $q = 1.5$). Same curves for spectral and TM analyses
293for the Parsivel² data are shown in Fig. 2.c and 2.d respectively. It appears that very similar
294results are found for the Parsivel² data. Similar curves for both devices are also obtained for
295the other studied quantity ($\rho, K_{dp}, Z_h, N_t, D_m$), therefore in the following only the curves for
296PWS100 will be showed and discussed in this sub-section.

297 Very similar results are found for ρ and K_{dp} (Fig. 3) with a very good scaling on the
298whole range of available scales on both the spectra and the TM analysis, which also has to be
299conducted on the fluctuations of the field. The good scaling behaviour retrieved for K_{dp} is not
300surprising and was actually expected if relation 17 is correct. Indeed a power (K_{dp} here) of a
301multifractal field (R here) should also behave as a multifractal field (this is the basic concept
302behind the DTM technique). To the knowledge of the authors, the multifractal behaviour of
303 K_{dp} has not yet been studied in time with disdrometer data this way, and it opens new
304perspectives. A potential one would be to compare these outputs with similar analysis
305performed in space with data provided by weather data. This would enable to study scaling
306relations in a spatio-temporal framework with a quantity directly measured by the radar (so
307far the only device providing “rather” high resolution space-time data of rain related fields)

308without having to rely on tailored relations that may introduce biases in the scaling behaviour
309as it is the case for the rain rate.

310 The situation for Z_h is more complex. Indeed the energy spectrum (Fig. 4.a) is not
311linear and could be interpreted as exhibiting two breaks, one at roughly 6 min^{-1} and the other
312one (the minimum on the curve) at roughly 2 min^{-1} . Authors do not have explanation for this
313behaviour which is not retrieved on R and K_{dp} . It was not expected and suggests that the
314“climatic” relation 16 does not hold very well. Indeed if it was true, a good scaling behaviour
315would be observed on the spectra. To confirm this, the same analysis were carried out on the
316quantity aR^b , where R is the rain rate studied before and a and b are the “climatic” values
317estimated in section 3.1. As expected for this analysis a good scaling behaviour is indeed
318retrieved on both the spectra (Fig. 4.c) and the TM analysis (Fig. 4.d). It should be mentioned
319that the greater scattering of the points for the $Z_h - R$ than the $R - K_{dp}$ relation observed on
320Fig. 1 is consistent with the fact that we found a good scaling behaviour on R and K_{dp} and not
321 Z_h . Quite surprisingly the scaling breaks observed on the spectra are not visible on the TM
322analysis (Fig. 4.b) where a unique regime is observed. It was implemented on the fluctuations
323of the field (Eq. 15) for which there are a slightly better scaling than for the field itself.
324However given the lack of scaling observed in the spectral analysis, the TM analysis might be
325not very reliable and should not be over-interpreted.

326 Finally we analysed the scaling features of N_t and D_m which are used to characterize
327the shape of the DSD. For N_t we find a good spectral behaviour but limited to the range 2 min
328– 2 h, as a flattening of the spectra is observed for small scales (Fig. 5.a). Similarly to the
329previously studied fields, a spectral slope greater than one is found meaning that the TM
330analysis should not be performed on the field directly but on its fluctuations (equation 15). It
331is displayed in Fig. 5.b and it appears that a good scaling behaviour is retrieved on a range of
332scales (4 min – 2 h) similar to the one observed on the spectra (Fig. 5.a). Considering the

333 whole range of scales would lead to a r^2 for $q = 1.5$ equal to 0.96 whereas is of 0.99 on the
334 limited range of scales. Similar scaling regimes are observed on D_m 's spectra (Fig. 6.a), but
335 the slope is much lower meaning that the TM analysis should be conducted directly on the
336 field. It yields a good scaling behaviour with a r^2 for $q = 1.5$ equal to 0.99 (it is of 0.87 if the
337 TM analysis is conducted on the fluctuations) (Fig. 6.b). The flattening of the spectra (almost
338 a horizontal slope) for small scales corresponds to what would be observed for a white noise.
339 This would suggest that N_t and D_m exhibit a scaling structure down to few minutes and behave
340 as a random homogeneous variable for smaller scales. It is not possible to confirm this
341 interpretation with the help of this data. More data at higher resolution, which would extend
342 the range of available small scales, would be needed to achieve this. A possible interpretation
343 of this could simply be that the sampling uncertainty is more visible at smaller scales and for
344 these quantities than the other ones.

345

346 3.3) UM parameters

347

348 Estimates of UM parameters for the various studied fields and the two devices are
349 reported in Table 2. It is timely to mention that the proportion of zeros is quite low (12 %,
350 with a fractal dimension of 0.96), which means that estimates are not biased by the
351 multifractal phase transition for small moment orders which is associated with them (see
352 Gires et al., 2012 for a in-depth analysis of this effect). Given the low quality of the scaling in
353 the spectral analysis for Z_h , the values of β and H are not shown for this field since they are
354 not reliable.

355 For all the fields we find H values greater than 0.5, except for D_m , which confirms that
356 the studied fields are non-conservative, and that the UM parameters α and C_1 should indeed
357 be estimated on their fluctuations as it was done. The estimates of the characteristic scaling

358 parameters are very similar for the two devices. It means that despite a roughly 10%
359 difference in terms of rain rate; they both record the same variability across scales. The
360 estimates for the rain rate, which is the field that has been mostly studied, are comparable
361 with the ones usually found for this range of scales by authors focusing the analysis on the
362 rainy portions (de Montera et al. 2009, Mandapaka et al. 2009, Verrier et al. 2010, Gires et al.
363 2013), although with slightly smaller values of α which were commonly reported to 1.8. With
364 regards to D_m and Z_h the differences between the two devices are more pronounced but it is
365 harder to interpret this fact given that the lower quality of the scaling for these fields implies
366 less reliable estimates.

367 It is possible to give an insight into radar relations 16 and 17 through the UM
368 parameters estimates. Indeed if a field is multifractal, then as previously mentioned a power of
369 it is also multiractal, and there is furthermore a relation between the α and C_1 (Tessier et al.
370 1993, Lovejoy et al. 2008) More precisely, if $R = cK_{dp}^d$, we have:

$$371 \alpha_R = \alpha_{K_{dp}} \quad (18.a)$$

$$372 C_{1,R} = d^\alpha C_{1,K_{dp}} \quad (18.b)$$

373 Here for both devices we find a roughly 0.2 difference between α computed for the two fields.

374 With regards to C_1 , for the PWS100 data $C_{1,R} / d_{PWS}^{\alpha_R} = 0.29$ (the value is slightly smaller

375 when considering $\alpha_{K_{dp}}$ in the relation) which is in rather good agreement with the 0.23

376 retrieved on the K_{dp} field. For the Parsivel² data $C_{1,R} / d_{Pars}^\alpha = 0.35$ which yields a value

377 greater than the 0.25 retrieved on the K_{dp} field. The same computations can be carried out for

378 the $Z_h - R$ relation. Indeed if $Z_h = aR^b$, one expects:

$$379 \alpha_{Z_h} = \alpha_R \quad (19.a)$$

380 $C_{1,Z_h} = b^\alpha C_{1,R}$ (19.b)

381 The difficulty here is that the agreement between the estimates of α is worse than for the
 382 $R - K_{dp}$ relation and the scaling on Z_h is not very good and reliable. With regards to C_1 we

383 have $b_{PWS}^{\alpha_{PWS,R}} C_{1,R_{PWS}} = 0.32$ (a slightly greater value is found using α_{Z_h}) which is quite

384 close to 0.34 found on Z_h , and $b_{Pars}^{\alpha_{Pars,R}} C_{1,R_{Pars}} = 0.39$ (a slightly greater value is found

385 using α_{Z_h}) which is quite different from the 0.51 found on Z_h . Finally it should be mentioned

386 that the estimates of α and C_1 on the field $a_{PWS} R_{PWS}^{b_{PWS}}$ (see Fig. 4.c and 4.d for scaling

387 curves) are respectively 1.55 and 0.28 which is roughly in agreement with equation 19 (

388 $b_{PWS}^{\alpha_{PWS}} C_{1,R_{PWS}} = 0.32$). The differences noticed with regards to the underlying theoretical

389 framework highlight the limitation of the “climatic” relations 16 and 17. The better agreement

390 in equations 18 and 19 for the PWS100 data is consistent with the greater quality of the fitting

391 of relations 16 and 17 for the PWS100 than for the Parsivel² (Fig. 1). The fact the α are equal

392 for R , ρ and N_t also suggests that a power-law relation between these quantities could be

393 investigated, which is not surprising since these quantities correspond to various moments of

394 the DSD (Sempere-Torres et al. 2000, Lee et al. 2004, Lovejoy and Schertzer 2008, Verrier et

395 al. 2013), even more directly than Z_h and K_{dp} , but this is beyond the scope of this paper.

396

397 4) Conclusions

398

399 In this paper the output data provided by two optical disdrometers is analysed; the

400 Campbell Scientific PWS100 based on the analysis of the light refracted by drops and the

401OTT Parsivel² based on the analysis of light occluded by drops. Not only the rainfall rate, but
402also the DSD parameters N_t and D_m , and the radar parameters Z_h and K_{dp} are studied.

403Furthermore the analysis is not performed only at the maximum resolution as it is commonly
404done but across scales with the help of the theoretical framework of Universal Multifractals.

405 It appears that despite a roughly 10 % differences in terms of rain rate, the multifractal
406analysis yield very similar results for the two measuring devices, with a slightly worse scaling
407observed on the Parsivel² data. R exhibits a very good scaling on the whole range of available
408scales (30 s – 2h) with $H=0.7$, $C_1=0.2$ and $\alpha=1.5$ which confirms findings of previous studies.
409A very good scaling is also retrieve on K_{dp} and UM parameters estimates are in rather good
410agreement with what would be expected if the standard relation $R = cK_{dp}^d$ is implemented
411with “climatic” parameters computed for this data set. The scaling of Z_h is worse, especially
412on the spectra, which highlights some limitations of the power law relation $Z_h = aR^b$. Finally
413the scaling behaviour only holds on the range few min – 2 h for N_t and D_m , and there are some
414hints at a possible random uniform behaviour for smaller scales possibly associated with
415sampling uncertainty.

416 These results suggest new ways of comparing the outputs of disdrometers by using
417other fields than the rain rate, and also scaling analysis. The results are particularly promising
418for K_{dp} which is also directly measured by polarimetric weather radars, contrarily to the rain
419rate for which non-trivial transformations potentially biasing the observed scaling are
420implemented. Multifractal investigations in a spatio-temporal framework on K_{dp} radar data
421should be carried out to improve knowledge about rainfall as a space-time process. More data,
422including spatial ones, should be analysed to confirm the possibility of characterizing DSD
423parameters with the help of UM, and hence the possibility of developing coupled multifractal
424cascades to actually simulate DSD fields.

425

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427

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432 Project (www.raingain.eu), and EU Climate KIC Blue Green Dream (www.bgd.org.uk)

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537

538

539

540**Table:**

541

| | | PWS100 | Parsivel ² |
|--------------|---|--------|-----------------------|
| $Z_h - R$ | a | 349 | 313 |
| | b | 1.49 | 1.63 |
| $R - K_{dp}$ | c | 17.3 | 15.3 |
| | d | 0.72 | 0.66 |

542Table 1: Parameters computed for the $Z_h - R$ relation (Equation 16) and the $R - K_{dp}$

543relation (Equation 17)

544

| | | R | ρ | K_{dp} | Z_h | N_t | D_m |
|-----------------------|----------|------|--------|----------|-------|-------|-------|
| PWS100 | β | 2.05 | 2.08 | 1.88 | - | 2.07 | 1.41 |
| | α | 1.51 | 1.53 | 1.35 | 1.72 | 1.58 | 1.66 |
| | C_1 | 0.18 | 0.16 | 0.23 | 0.34 | 0.11 | 0.005 |
| | H | 0.67 | 0.67 | 0.62 | - | 0.63 | 0.21 |
| Parsivel ² | β | 1.96 | 2.08 | 1.71 | - | 2.11 | 1.33 |
| | α | 1.57 | 1.54 | 1.33 | 2.00 | 1.60 | 2.10 |
| | C_1 | 0.18 | 0.16 | 0.25 | 0.51 | 0.11 | 0.006 |
| | H | 0.63 | 0.67 | 0.55 | - | 0.64 | 0.17 |

545 Table 2: Scaling parameters of the various studied fields for the PWS100 and Parsivel² data

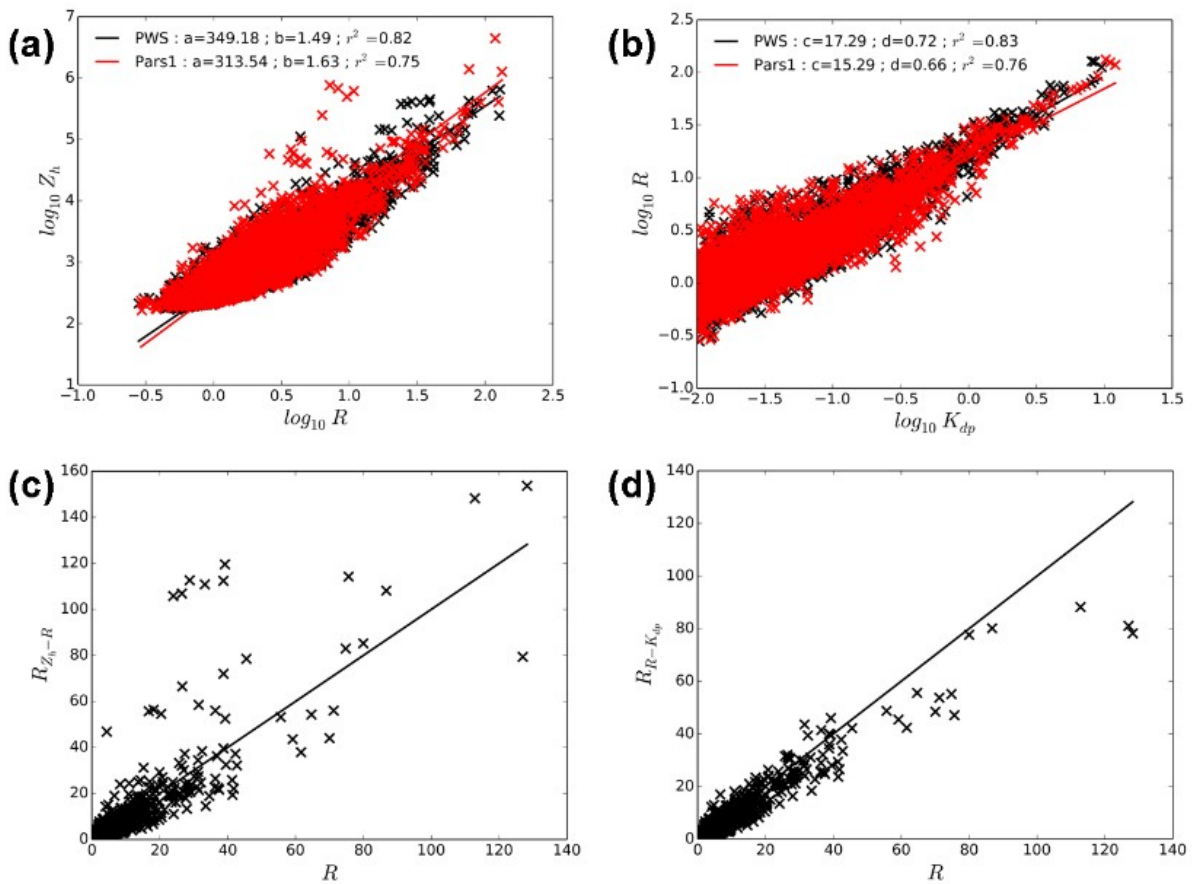
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549 **Figure caption list:**

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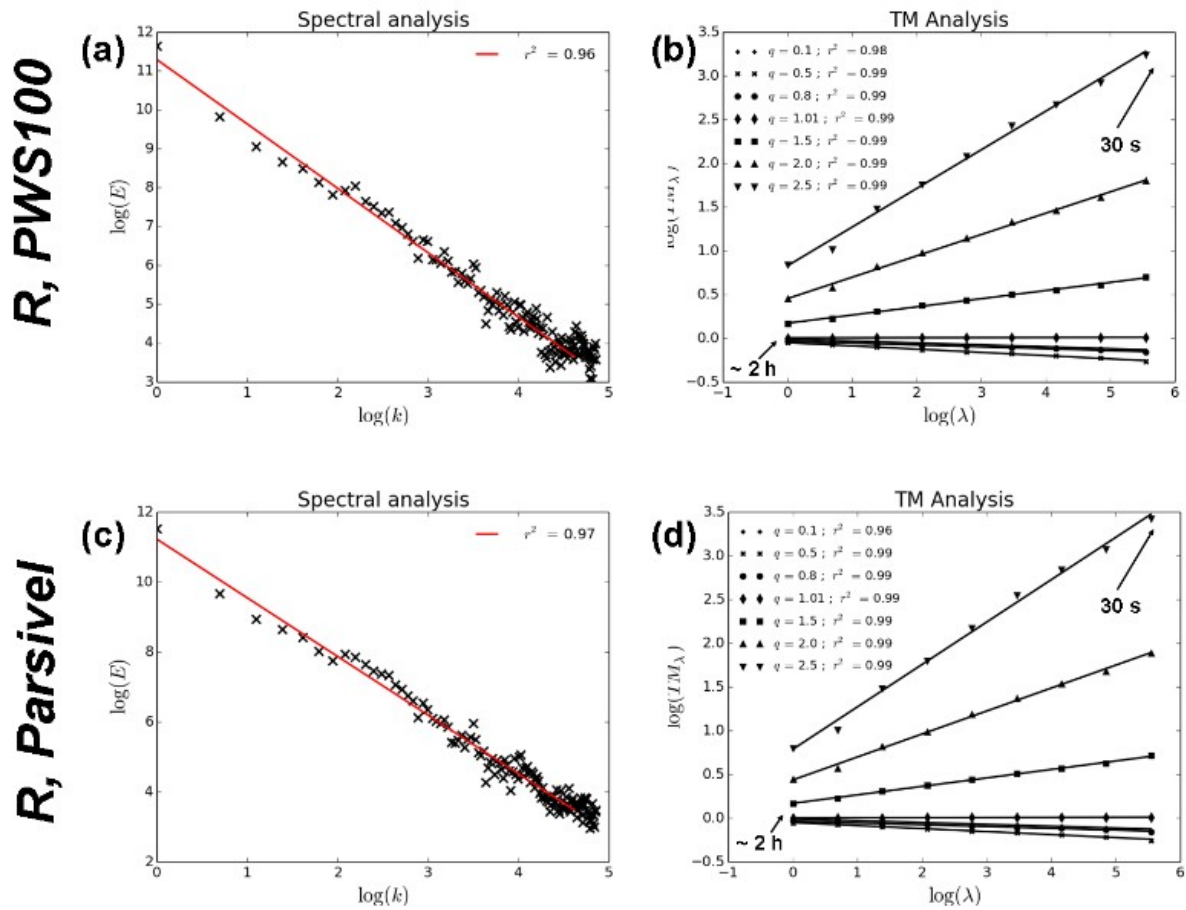
551

552 Figure 1: Computation of the parameters of the $Z_h - R$ (a) and $R - K_{dp}$ (b) relations

553 (equations 16 and 17 respectively in log-log plot). Scatter plot of R_{Z-R} (c) and $R_{R-K_{dp}}$ (d)

554 versus R for the PWS100 data

555



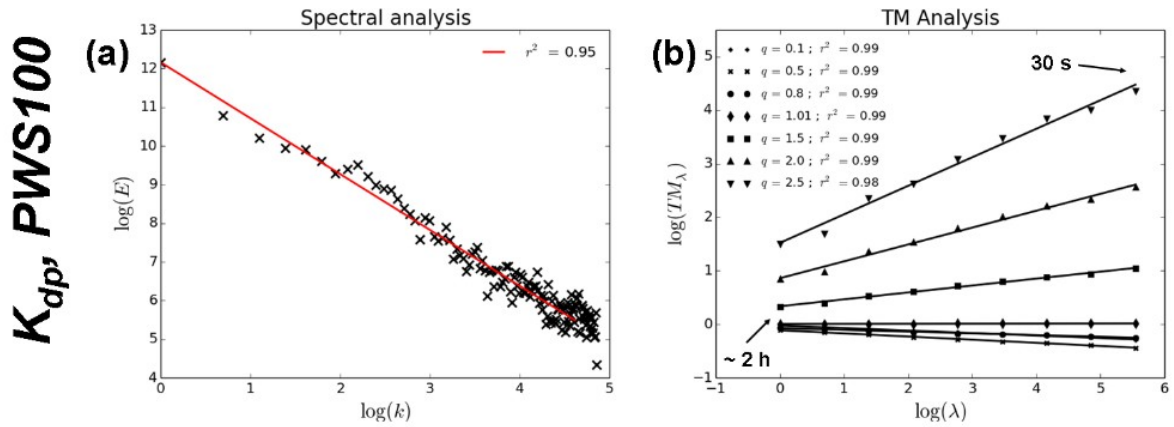
556

557 Figure 2: Scaling analysis for R measured by the PWS100: (a) Spectral analysis, i.e. equation

558 14 in log-log plot; (b) Trace Moment (TM) analysis, i.e. equation 9 in log-log plot. (c) and (d)

559 Same as (a) and (b) for R measured by the Parsivel².

560

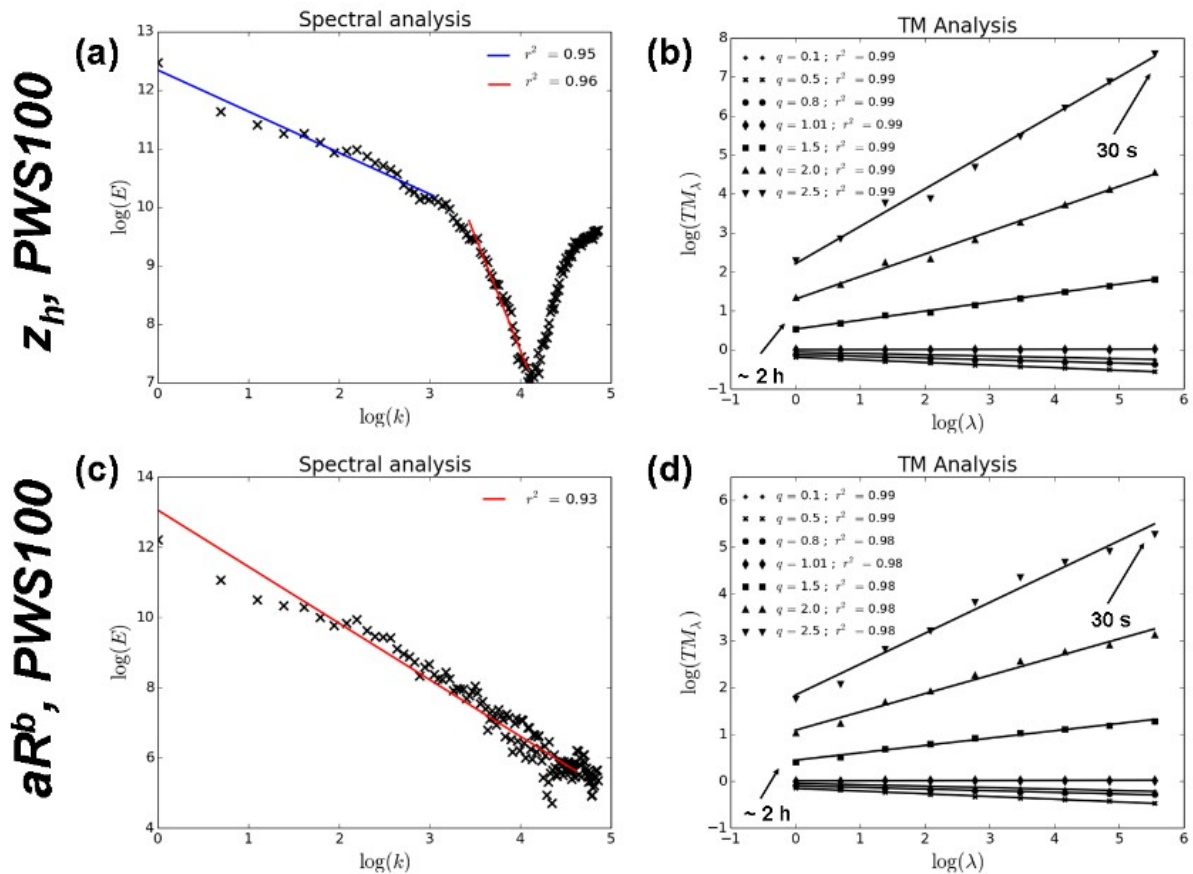


561

562 Figure 3: Scaling analysis for K_{dp} measured by the PWS100: (a) Spectral analysis, i.e.

563 equation 14 in log-log plot; (b) Trace Moment (TM) analysis, i.e. equation 9 in log-log plot

564



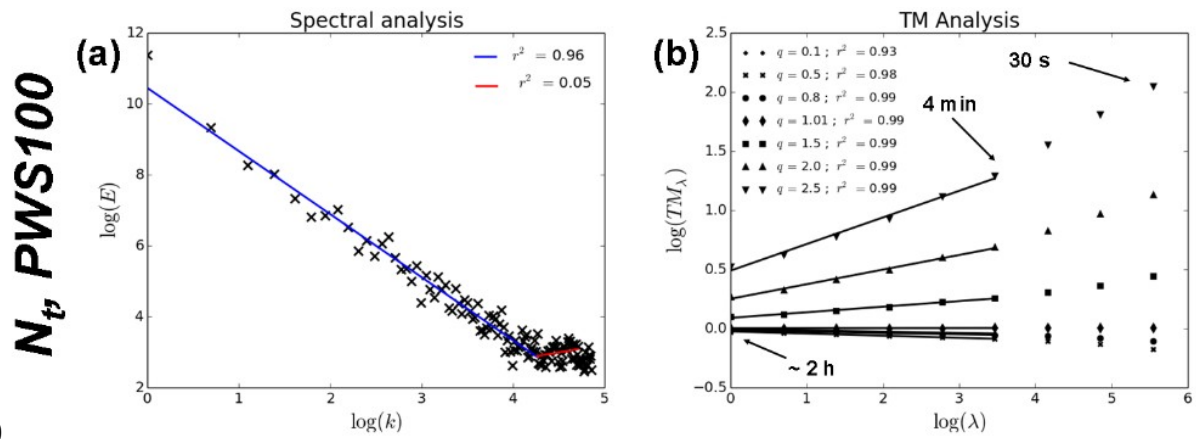
565

566 Figure 4: (a) – (b) Same as in Fig. 3 but for Z_h measured by the PWS100, in (a) a break is

567 considered for $k = 20-30$ ($\sim 6 \text{ min}^{-1}$) and $k = 60$ ($\sim 2 \text{ min}^{-1}$); (c) – (d) Same as in Fig. 3 but for

568 aR^b measured by the PWS100 and “climatic” parameters a and b found in section 2.

569

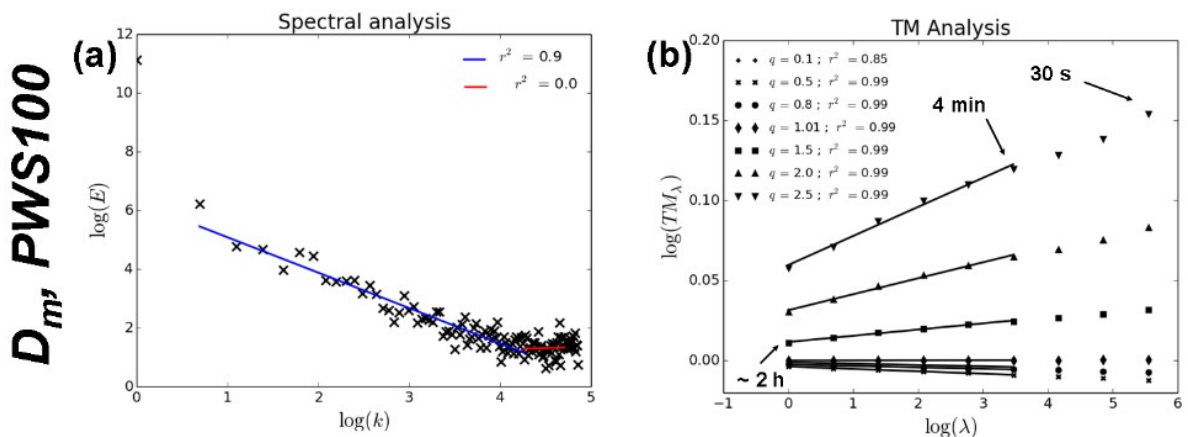


570

571 Figure 5: Same as in Fig. 3 but for N_t measured by the PWS100, in (a) a break is considered

572 for $k=70$ ($\sim 2 \text{ min}^{-1}$), in (b) a break is considered for $\lambda=32$ (4 min)

573



574

575 Figure 6: Same as in Fig. 3 but for D_m measured by the PWS100, in (a) a break is considered

576 for $k=70$ ($\sim 2 \text{ min}^{-1}$), in (b) a break is considered for $\lambda=32$ (4 min)

577